

Last week =

H atom spectrum -

Lyman, Balmer etc. series.  
Experiment is measuring a quantity of a discharges gas -

$$\tilde{\nu} \sim R \left( \frac{1}{m^2} - \frac{1}{n^2} \right)$$

m, n integers -

$\tilde{\nu}(+) \Rightarrow n > m$

via this you can observe any frequency in H.  
Every line you can observe is given by this formula.

Rydberg combinatory principle =

m = constant

change n and subtract from  
each other (the  $\tilde{\nu}$ 's are also a line in  
the spectrum.

JJ Thompson try to manipulate the cathode rays  
with  $\vec{E}$  field and  $\vec{B}$  field. What is the result.

Slide = "Elektro katodemas"

In the experiment particles are "sagittok"

Go through the theory JJ went through (like mass ions  
etc.) while discovering electron.

Also check the oil drop exp.

RUTHERFORD EXPERIMENT = Planetary motion of  
particles

1903

H. Nagaoke -

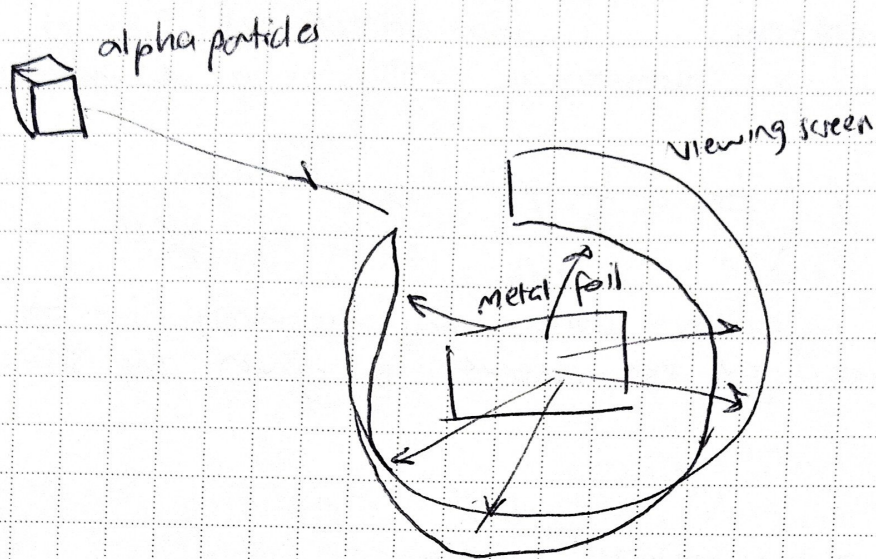
planetary model

1906

E. Rutherford -

starts testing and confirms  
in 1911

First particle physics experiment  
actually.



Type of materials when an energetic particle heats them, they splash photons.  
 What are they called  
 Subs. - s... something.

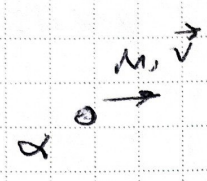
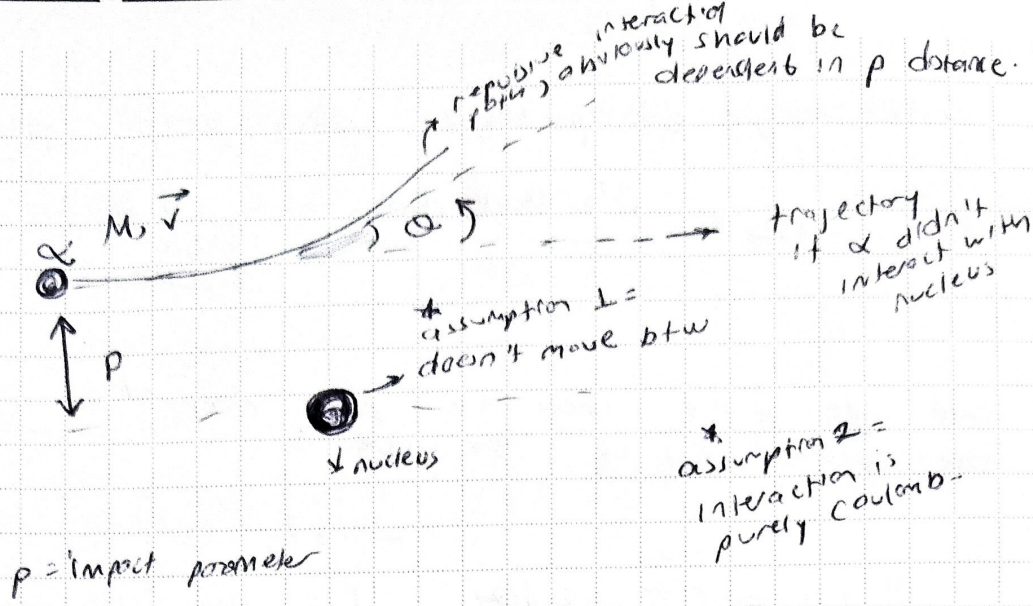
$\alpha$  particle  ${}^4_2\text{He}$

$\alpha$  particle has identical energy from radioactive nuclei!!!  
 Well-defined energy.

${}^4_2\text{He}$  is heavy, but Gold nuclei is also big.

The experiment compares two scenarios =

a uniform sheet of positive charges vs. some nuclei with planetary-like motion forming the sheet.



$$F = \frac{2Ze^2}{4\pi\epsilon_0 r^2}$$

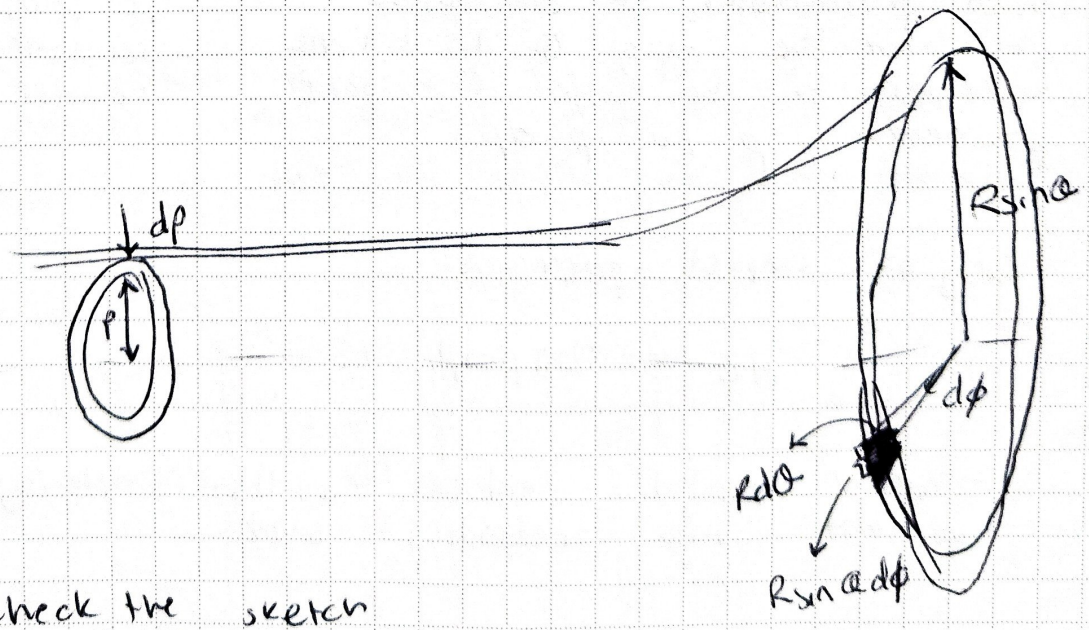
Now consider "Kepler problem" in order to derive formulation for trajectory of  $\alpha$ .

$$d\theta \frac{Q}{2} = 2\pi\epsilon_0 \frac{Mv^2 p}{Ze^2}$$

DERIVE THIS!!!!

is this cotangent?

We would like a distribution of angles - with parameters of the alpha particle.



check the sketch from slides!

Solid angle of particles with  $p \in p$ ,  $p \in dp$

$$d\Omega = \frac{R^2 \sin\theta d\theta d\phi}{R^2}$$

solid angle  
ELEMENT

solid angle is the area on a sphere compared to the whole area of the sphere.

$$d\Omega = 2\pi \sin\theta d\theta \quad (\text{is this still the element? } \Omega \text{ or } d\Omega)$$

Probability of scattering at  $\theta$ ,  $\theta + d\theta$

$$\frac{dN}{N} = l \cdot n \cdot d\Omega$$

$\frac{dN}{N}$  → particles for a particular angle  
 $N$  → tot num of particles

differential cross section

why a cross section?  
because of units!  
check it out!

$$\frac{dI}{I} = \Sigma c l$$

nucleus is very likely to scatter at this angle is high!  
if  $\Omega$  is high?  
if low, scattering likelihood is also low!

! When  $p$  increases,  $\theta$  decreases, you have the ring  $\theta$  to  $\theta + d\theta$  corresponds to the range of impact parameters  $p$  to  $p + dp$ .

Ring of impact parameters:

$$d\Omega = 2\pi p dp$$

thinner so that metal this reduces statistics multiple scattering work.

if we only count single count collisions!

$$dN_{\alpha} = \rho \cdot n \cdot N \cdot d\Omega_{\alpha}$$

no double scattering events.

Trajectory -

$$\rho^2 = \left( \frac{ze^2}{2\pi\epsilon_0 M v^2} \right)^2 \cdot \cot^2 \frac{\alpha}{2} \quad \text{differentiate.}$$

↓  
kotanjent.  
 $\left( \frac{\cos ?}{\sin ?} \right)$  idk.

$$p d\rho = \left( \frac{ze^2}{2\pi\epsilon_0 M v^2} \right)^2 \frac{\cot^2 \frac{\alpha}{2}}{\sin^2 \frac{\alpha}{2}} d\left(\frac{\alpha}{2}\right)$$

$$d\alpha = \frac{d\Omega}{2\pi \sin \alpha} = \frac{d\Omega}{4\pi \sin \frac{\alpha}{2} \cos \frac{\alpha}{2}}$$

plug this into the probability of scattering ( $\rho \cdot n \cdot d\Omega_{\alpha}$ )

$$dN_{\alpha} = \rho \cdot n \cdot N \left( \frac{ze^2}{4\pi\epsilon_0 M v^2} \right)^2 \frac{d\Omega}{\sin^4 \frac{\alpha}{2}} \quad *$$

↙ notice double KE

\* Rutherford Formula for scattering probability!!!

You obtain a constant from this expression via something idk what or how tho.

Also compare the  $\alpha - dN_{\alpha}$  relations. Analyse.

### RESULT

| $\alpha, \text{deg}$ | $dN_{\alpha}$ | $dN_{\alpha} \cdot \sin^4 \frac{\alpha}{2}$ |
|----------------------|---------------|---|
| 15                   | 13200         | 38.4  |
| 30                   | 7800          | 35  |
| 45                   | 4435          | 20.8  |
| 60                   | 477           | 23.8  |
| 105                  | 69            | 27.5  |
| 150                  | 33            | 28.5  |

check the  $dN \sin^4(\alpha)$  vs.  $\alpha$  graph from slides as well.

very important!

$$\frac{dN}{N} = \ln 6$$

↳ cross section of nucleus.

You can estimate nucleus radius as well...

This was Rutherford type of model.

One problem with planetary-like model of atoms is leading to the use of Maxwell. Maxwell tells, spinning of electron around nuclei should be forming an E field.

Therefore (due to centripetal force so an acceleration) an B field. So atom should be emitting. But it doesn't...

There's also uncertainty as well.

Around 1913's, it is started to be discussed that atoms have levels, discrete and well-defined.

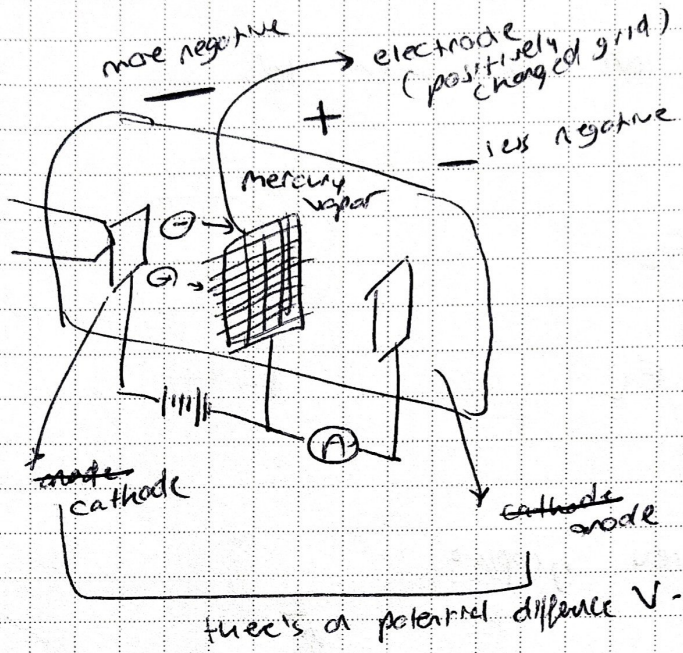
# Frank-Hertz Experiment.

Assumption =

$e^-$  in atoms have discrete energies.

$E_1, E_2, E_3 \dots$  etc.

Frank-Hertz tests this.



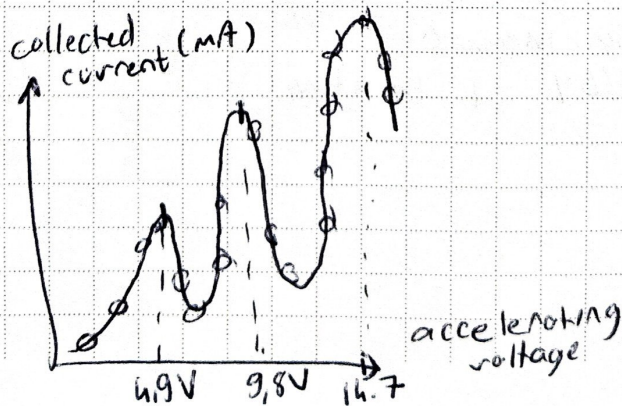
1st excitation =  $E_2 - E_1$

as we have enough energy satisfied to bring from lower to higher level

Increase voltage  $\uparrow$   
Current  $\uparrow$

HOWEVER,

! when you reached the energy between the electrode and the less negative anode, electron goes to the electrode.  
\* So number of electrons reaching DROP as excitation occurs!!!



Frank-Hertz Data for Mercury.

4.9V is the dropping point.

At dropping points, there's a collisions of electrons between anode and grid of some energies. Electrons end up going to the positive grid.

We indeed proved, electrons have discrete energy.

\*  $4.9 \text{ eV}$  is the first excitation energy for Mercury.

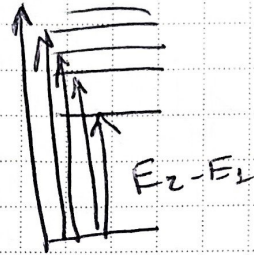
What further we can do with this experiment?

We have

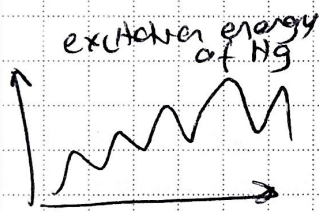
$$E_2 - E_1$$

but also,

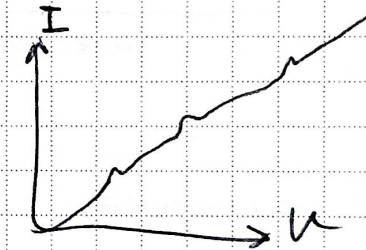
$$E_3 - E_1$$



Now with the prev graph =



Less atoms



signals less.  
BUT, higher probability to obtain all excitations.

Reduce pressure = measure

$$E_2 - E_1$$

$$E_3 - E_1$$

Now we can also measure IONIZATION ENERGY of the atom

$$E_{\text{ionization}} = 13.6 \text{ eV} \quad (\text{H-atom})$$

N Bohr's model of atom

Postulates =

i)  $e^-$ 's are in fixed orbits with energies  $E_1, E_2, E_3, \dots$ . No absorption or emission takes place when orbiting.

ii) in the H atom,  $e^-$  angular momentum is

$$L = n\hbar, \quad n = 1, 2, 3, \dots$$

Planck =  $E_n = n\hbar\omega$        $[\hbar] = \text{J}\cdot\text{s}$  (dim of angular momentum)

Bohr =  $L = n\hbar$

When changing orbits, atom will emit or absorb light (related to energy difference) of the orbitals.

(iii) Atom radiates/absorbs light when changing orbits and the frequency is =

$$\hbar\omega_{mn} = E_n - E_m$$

In a circular orbit =

$$F = \frac{ze^2}{4\pi\epsilon_0 r^2} \quad a = \frac{v^2}{r}$$

II Newton Law

$$F = ma \quad \frac{mv^2}{r} = \frac{ze^2}{4\pi\epsilon_0 r^2}$$

Full energy of electron =

$$E = T + U$$

$$T = \frac{1}{2} mv^2$$

$$U = -\frac{ze^2}{4\pi\epsilon_0 r}$$

bound (-) total energy

$$F_{\text{tot}} = \frac{mv^2}{r} - \frac{ze^2}{4\pi\epsilon_0 r}$$

substitute =

$$E_{\text{tot}} = \frac{-ze^2}{8\pi\epsilon_0 r} \quad \frac{ze^2}{4\pi\epsilon_0 r^2}$$

Now we will investigate how far  $e^-$  can be from the nucleus.

2nd postulate =

in circular motion =

$$p = mv \Rightarrow I\omega$$

$$L = I\omega = mr^2\omega$$
$$= n\hbar; \quad \omega = \frac{v}{r}$$

$$L = mr^2\omega = n\hbar = mvr$$

$$\frac{mv^2}{r} = \frac{r m v^2}{r^2} = \frac{ze^2}{4\pi\epsilon_0 r^2}$$

$$\frac{1}{r^2} \frac{1}{m} \underbrace{m^2 v^2 r^2}_{L^2} = \frac{ze^2}{4\pi\epsilon_0 r} = \frac{1}{r^2} \frac{1}{m} n^2 \hbar^2$$

radius of the orbit ranked  $n =$

$$r_n = \frac{4\pi\epsilon_0 n^2 \hbar^2}{m z e^2} = \underbrace{4\pi\epsilon_0 \frac{\hbar^2}{m z e^2}}_{\text{unit is a length}} n^2$$

unit is a length.

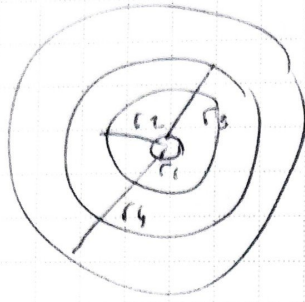
It is the Bohr radius!!!

BOHR RADIUS

$$r_B = 4\pi\epsilon_0 \frac{\hbar^2}{m e^2}$$

0,53 Å

$$r_n = r_B \cdot n^2 \quad (z=1)$$



$$\begin{aligned} r_1 &= r_1 \\ r_2 &= 4r_1 \\ r_3 &= 9r_1 \\ &\vdots \\ r_4 &= 16r_1 \end{aligned}$$

$$E_n = -\frac{ze^2}{8\pi\epsilon_0 r_n}$$

$$E_n = -\frac{mz^2e^4}{(4\pi\epsilon_0)^2 2\hbar^2} \cdot \frac{1}{n^2}$$

3rd Postulate =

$$h\nu_{nk} = hc\bar{\nu}_{nk} = z^2hc\bar{\nu}_{nk} = E_n - E_k$$

↓  
freq from orbit n to k

$$\bar{\nu}_{nk} = \frac{1}{(4\pi\epsilon_0)^2} \cdot \frac{mz^2e^4}{4\pi\hbar^3c} \left( \frac{1}{k^2} - \frac{1}{n^2} \right)$$

$$\nu = R \left( \frac{1}{m^2} - \frac{1}{n^2} \right) \text{ Balmer Formula!}$$

Bohr's end up done with calculating Rayleigh constant!!!  
or  
Reidburg idk

From Bohr =

$$R = \frac{1}{(4\pi\epsilon_0)^2} \frac{me^4}{4\pi h^3 c} \dots$$

$$R^{\text{theory}} = 109737,303 \text{ cm}^{-1}$$

$$R^{\text{exp}} = 109677,581 \text{ cm}^{-1}$$

$$\text{accuracy} = 6 \cdot 10^{-4}$$

Can I do better? Said Bohr.

The error was due to mass, gravity etc -

Consider motion of nucleus =  
reduced mass.

$$m' = \frac{m}{1 + \frac{m}{M}}$$

$m = e^-$  mass  
 $M = \text{nucleus mass}$

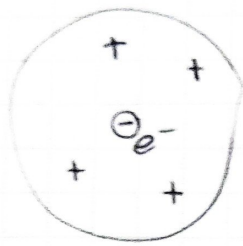
by replacing  $m_e \rightarrow m'$   
accuracy is  $\sim 10^{-7}$

For other nuclei = for other atoms ( $Z \neq 1$ ,  $A \neq H$ )

$$E_n = Z^2 \cdot (13.6 \text{ eV}) \frac{1}{n^2}$$

## EXERCISES

- I Find the radius of H atom given the ionization energy of 13.6 eV. Apply Thompson's pudding model.



### SOLUTION

electron - in the center of a charged ball

Inside el. field

$$E_1 = \frac{er}{4\pi\epsilon_0 R^3} \Rightarrow \text{Gauss Theorem.}$$

↓  
flux through the surface of hypothetical surface.

$$\begin{aligned}\Phi &= EA \\ &= \frac{1}{\epsilon_0} Q'\end{aligned}$$

outside el. field

$$E_2 = \frac{e}{4\pi\epsilon_0 R^2}$$

Now we know the force applied on electron.

$$W_{ionization} = \frac{e^2}{4\pi\epsilon_0} \left[ \int_0^R \frac{r}{R^3} dr + \int_R^\infty \frac{dr}{r^2} \right]$$

$$= \frac{e^2}{4\pi\epsilon_0} \left[ \frac{1}{R^3} \frac{r^2}{2} \Big|_0^R - \frac{1}{r} \Big|_R^\infty \right] =$$

$$\frac{e^2}{4\pi\epsilon_0} \left[ \frac{1}{2R} + \frac{1}{R} \right]$$

$$= \frac{e^2}{4\pi\epsilon_0} \frac{1}{R} \frac{3}{2} ; \text{Wronstein} = \frac{3}{8} \frac{e^2}{\pi\epsilon_0} \frac{1}{R}$$

$$R = \frac{3}{8} \frac{e^2}{\pi\epsilon_0} \frac{1}{Wron}$$

$$R = \frac{3}{8\pi} \cdot \frac{1.6 \times 10^{-19}}{8.85 \cdot 10^{-12}} \cdot \frac{1.6 \cdot 10^{-19}}{13.6 \text{ eV}}$$

to get Joules

$$= 1.587 \cdot 10^{-10} \text{ m} = 1.59 \text{ \AA}$$

radius of H from Thomson model.

II Estimate the number of particles will be scattered in angle range  $46^\circ$  to  $46^\circ$  if we fire  $10^4$  particles with the E of 1 MeV, into Cu (copper) foil with thickness  $l = 5 \mu\text{m}$ . Calculate the effective scattering cross section.

SOLUTION

$\frac{dN}{N} \rightarrow$  number of particles scattered for a certain angle range.

$$\frac{dN}{N} = \rho \cdot l \cdot \frac{Z^2 e^2}{4\pi\epsilon_0 M v^2} \frac{2\pi \sin^2 \frac{\theta}{2} d\theta}{\sin^4 \frac{\theta}{2}}$$

in this problem, due to small angle range, we can take  $45^\circ$  as a constant and simply multiply w/  $d\theta = 2^\circ$  and no integration needed. Normally you would tho.

$$\rho_{\text{cm}} = 8.9 \frac{\text{g}}{\text{cm}^3}$$

$$A = 64 \frac{\text{g}}{\text{mol}}$$

$$Z = 29$$

$$n = \frac{\rho N_A}{A}$$

how many nuclei per cubic meter

$$= \frac{8.9}{64} \cdot 6.022 \cdot 10^{23}$$

$$= 8.372 \cdot 10^{22} \text{ m}^{-3}$$

$$= 10^{22} \text{ cm}^{-3}$$

$$\alpha \approx 45^\circ$$

$$\sin \frac{45^\circ}{2} \approx 0.38$$

$$d\alpha = 2^\circ = \frac{2}{57} = 0.035$$

$$\frac{mv^2}{2} \text{ KE of } \alpha \text{ particle.} = 1 \text{ MeV} = 1.6 \times 10^{-13} \text{ J}$$

$$dN = 13.2 \text{ particles.}$$

⇒ cross section is a probability that is attenuation related (what?)

$$\frac{dN}{N} = \underbrace{L}_{\substack{\downarrow \\ \text{length} \\ \text{of the} \\ \text{foil, how} \\ \text{long the} \\ \text{scattering} \\ \text{thing is}}} \cdot \underbrace{n}_{\substack{\text{m}^{-3} \\ \text{therefore,} \\ \text{area}}} \cdot \underbrace{d\Omega}_{\substack{\text{likelihood} \\ \text{of scattering} \\ \text{at } \alpha}}$$

$\sigma$  ⇒ describes the interaction of particles, or radiation

Higher probability, higher the area of cross section

But cross section is not a probability since it has the measurements of area  $m^2$ .

Wtf is it then?

$$\frac{dN}{N} = f \cdot A \cdot d\Omega$$

$$d\Omega = \frac{13.2}{10^4 \cdot 5 \cdot 10^{-6} \cdot 8.37 \cdot 10^{28}} = 0.51 \cdot 10^{-26} \text{ m}^2$$

conversion =  $\text{m}^2 \rightarrow \text{cm}^2$

$$0.51 \cdot 10^{-22} \text{ cm}^2$$

$$3.1 \cdot 10^{-23} \text{ cm}^2$$

what is 1 barn?

$$1 \text{ barn} = 10^{-24} \text{ cm}^2$$

$$d\Omega = 31 \text{ barn}$$

III In H spectrum, two Balmer series lines have wavelengths  $\lambda_1 = 410.2 \text{ nm}$   
 $\lambda_2 = 486.1 \text{ nm}$

which series ( $m=?$ ) will contain the combination line?

SOLUTION

Balmer series =

$$\bar{\nu} = R \left( \frac{1}{2^2} - \frac{1}{n^2} \right) \quad \bar{\nu} = \frac{1}{\lambda}$$

$$\frac{1}{\lambda} = R \left( \frac{1}{4} - \frac{1}{n^2} \right); \quad \frac{1}{n^2} = \frac{1}{4} - \frac{1}{R\lambda} =$$

$$= \frac{R\lambda - 4}{4R\lambda} \Rightarrow n^2 = \frac{4R\lambda}{R\lambda - 4}$$

$$n_1 = \frac{4R\lambda_1}{R\lambda_1 - 4}, \quad n_2 = \frac{4R\lambda_2}{R\lambda_2 - 4}$$

$$= 6$$

$$= 4$$

combination line =

$$\bar{r} = R \left( \frac{1}{4^2} - \frac{1}{6^2} \right)$$

$\downarrow$   
 $m=4$  //

$\Rightarrow$

acknowledge,  
(Brockett)